

SM2 Micromegas Modules in the Cosmic Ray Facility

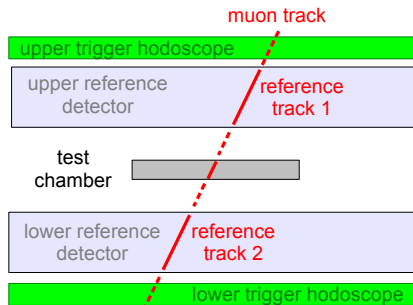
Maximilian Herrmann

Ludwig-Maximilians-Universität München - Lehrstuhl Schaile

Muon Week 20.10.2020

Quality Control Using Cosmic Muon Passages

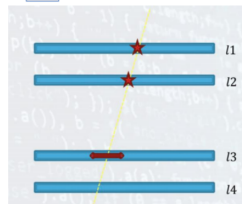
MDT tracks



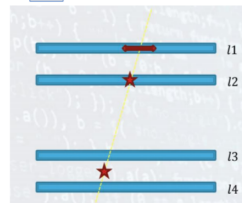
comparison of both methods
to check for compatibility

internal

(a)

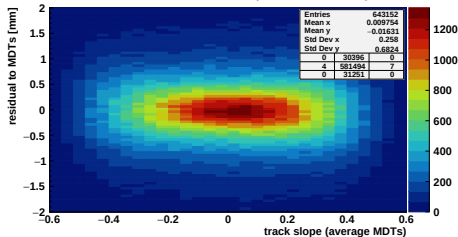


(b)

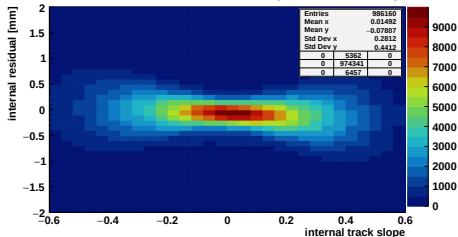


Comparison of External and Internal Tracking (e.g. M29)

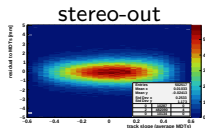
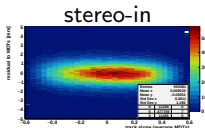
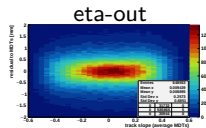
MDT tracks (at eta-in)



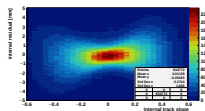
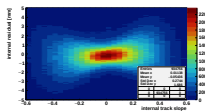
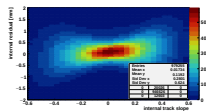
internal tracking (for eta-in)



MDT
tracks



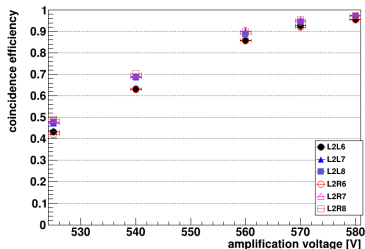
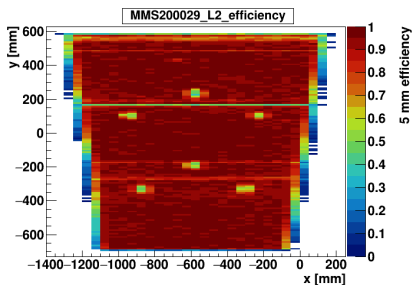
internal



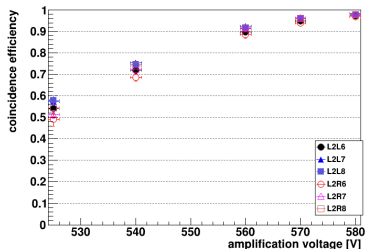
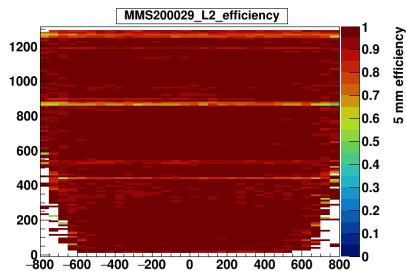
! other y-scale for stereos !

Comparison of External and Internal Efficiency Estimation (e.g. M29, eta-in)

MDT tracks

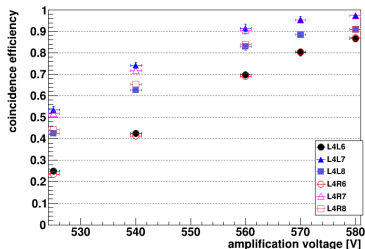
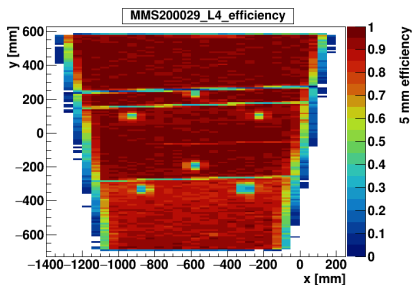


internal

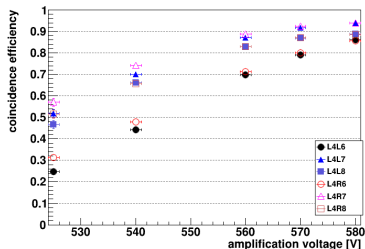
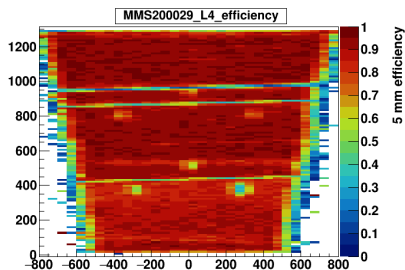


Comparison of External and Internal Efficiency Estimation (e.g. M29, stereo-out)

MDT tracks



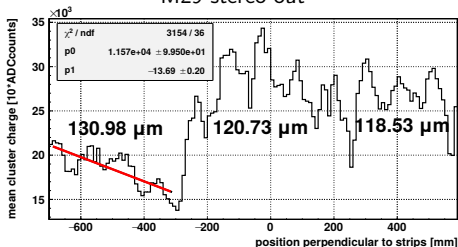
internal



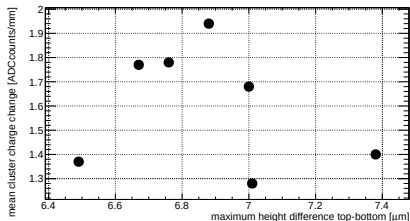
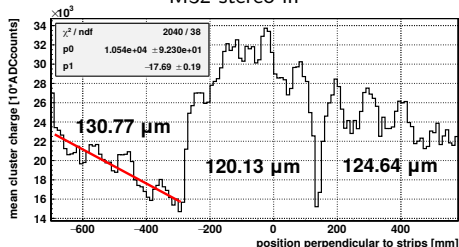
Effect of Pillarheight Trend on Cluster Charge (at 580V)

affected PCB of SS6-type $\Rightarrow$ stereo boards 6 of Modules 29 to 32

M29 stereo-out



M32 stereo-in

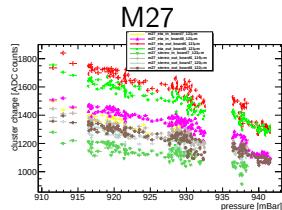
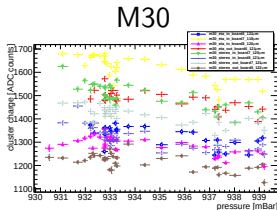
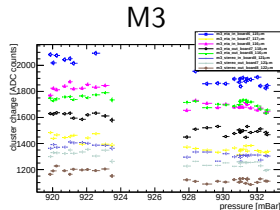
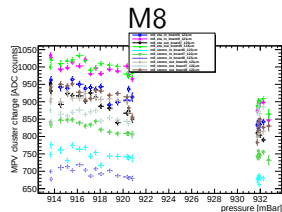
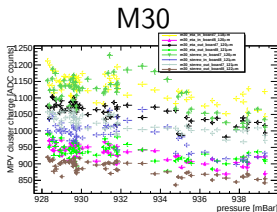
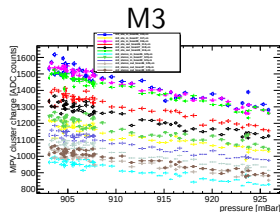


$\Rightarrow$ no clear correlation between height difference and cluster charge change

Pulseheight Dependencies - Pressure and Pillarheight

pressure dependence fitted for each readout-board individually

⇒ evaluation as function of pillarheight (linear fit)



data usable for modules :

570 V : 3, 4, 5, 6, 8, 9, 11, 12, 14, 15, 19, 20, 26, 30

580 V : 13 , 17 , 18 , 21 , 23 , 24 , 25 , 27

Pulseheight Dependencies - Module Summary

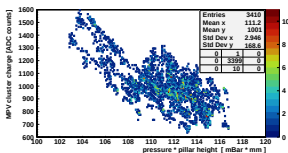
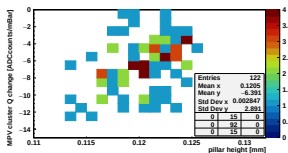
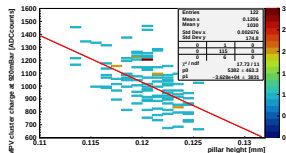
MPV cluster charge $\Rightarrow$

at 920 mBar *
VS pillar height

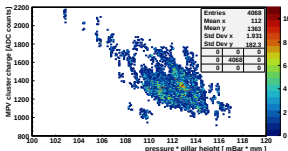
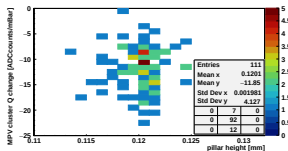
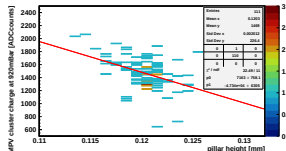
pressure dependence
(fit slope)

pillar height
times pressure

570 V



580 V



$\Rightarrow$ relative pulseheight change at 120 μm :

570 V : $-3.5 \pm 2.0 \text{ \%} / \mu\text{m}$

580 V : $-3.2 \pm 2.1 \text{ \%} / \mu\text{m}$

* pressure offset about -30 mBar

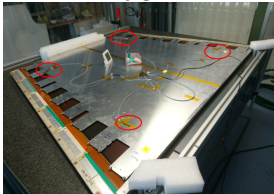
Alignment Analysis by Fabian Vogel (new PhD student)

results of his Masterthesis

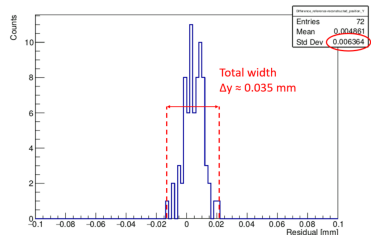
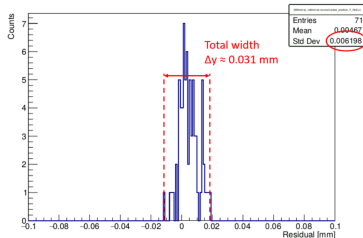
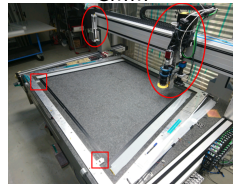
https://www.etp.physik.uni-muenchen.de/publications/theses/download/master_fvogel.pdf

⇒ calibration of jig and CMM using Saclay reference data (panel eta10, stereo11)

jig

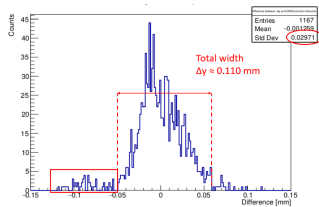


CMM

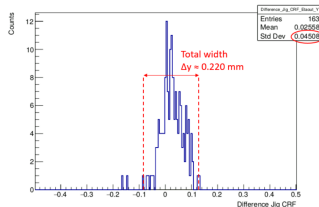


Compatibility Reconstruction of Measurements by Fabian Vogel

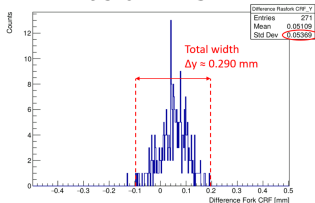
jig - CMM



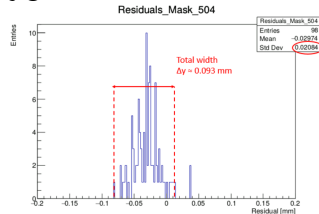
jig - CRF



rasfork - CRF



jig residuals for mask 504



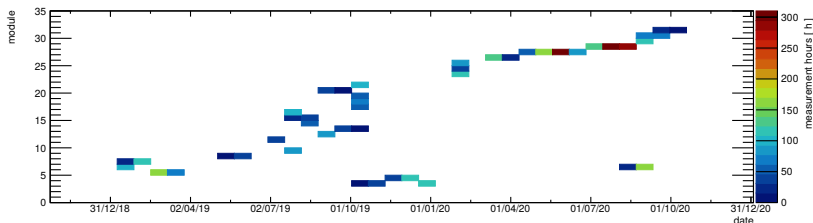
⇒ good agreement between measurements within error estimation

⇒ summary of single mask residuals reconstruct systematics during gluing

Summary

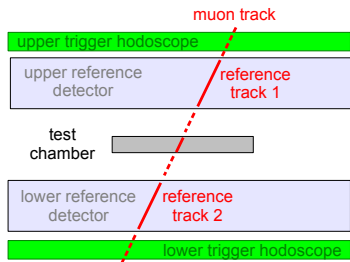
quality control of the full SM2 active area using cosmic muons

- alternative method for tracking implemented to check efficiency reconstruction
⇒ compatible results of internal and external tracking
- pulsheight dependencies investigated:
 - pillarheight-trends visible
⇒ magnitude not yet correlated
 - pressure-dependence for each PCB individual
⇒ pillarheight correlation in the order of -3% / μm
- improved alignment reconstruction by Fabian Vogel using multiple measurements
⇒ calibration of jig and CMM



Backup

Validation using Cosmic Muons - Cosmic Ray Facility



trigger

track reconstruction

active area

angular acceptance

energy cut

readout (full module)

readout rate

scintillator hodoscope

resolution ~ 10 cm

Monitored Drift Tubes (MDTs)

resolution ~ 0.2 mm

$2.2 \text{ m} \times 4 \text{ m}$

$\pm 30^\circ$

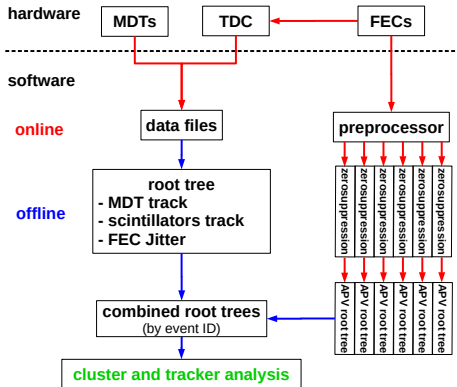
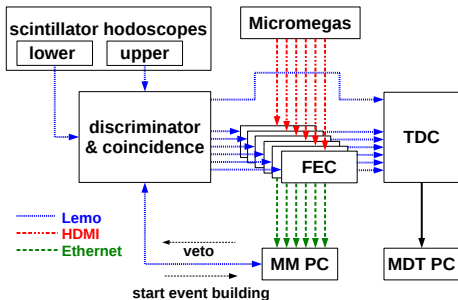
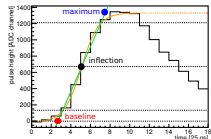
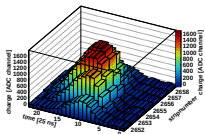
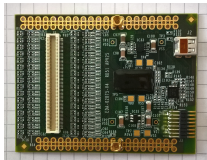
iron plate $\rightarrow E_\mu > 600 \text{ MeV}$

12288 channels

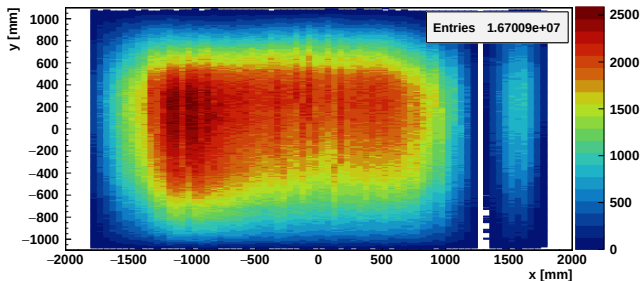
100 Hz (online zerosuppression)



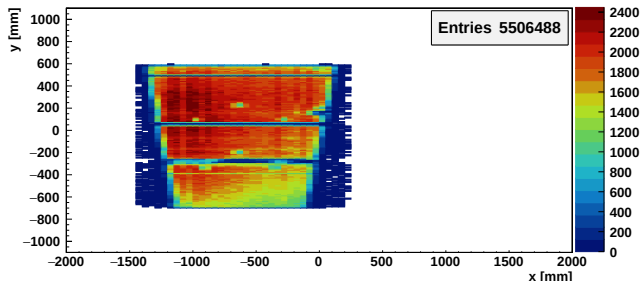
SRS-based Data-Acquisition



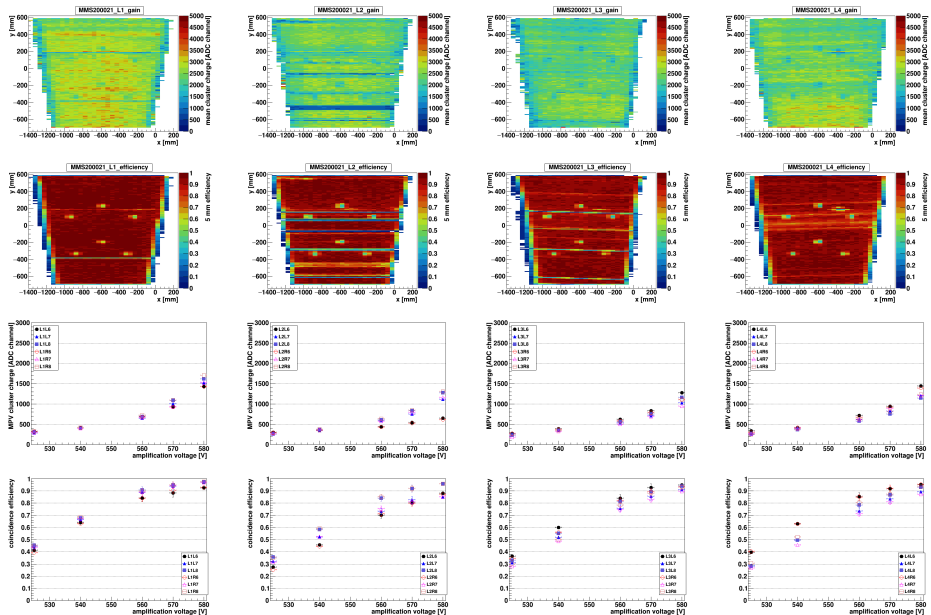
CRF hit distribution



M3 eta_out, 5 mm efficient

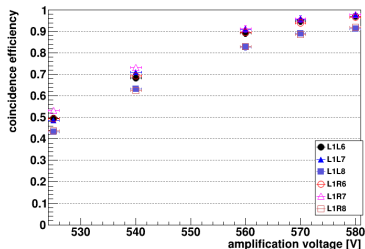
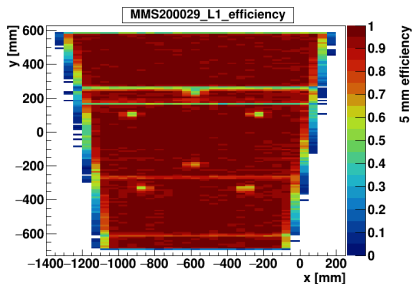


Pulse Height and Efficiency (Ar:CO₂ 93:7 vol%) Module 21

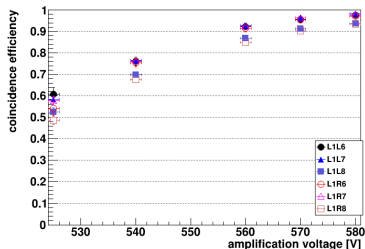
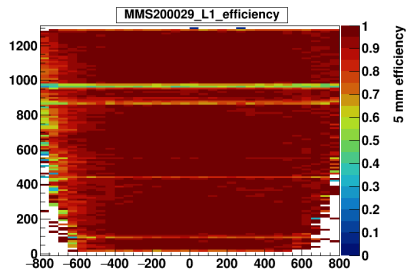


Comparison of External and Internal Efficiency Estimation (e.g. M29, eta-out)

MDT tracks

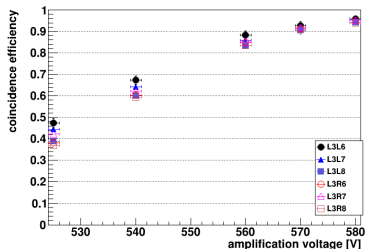
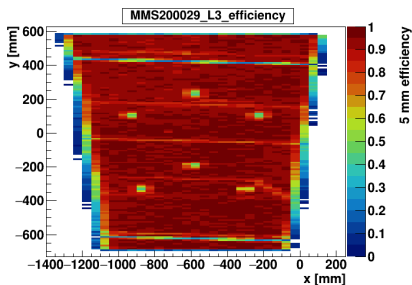


internal

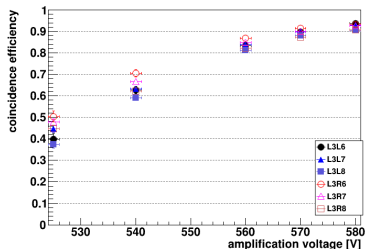
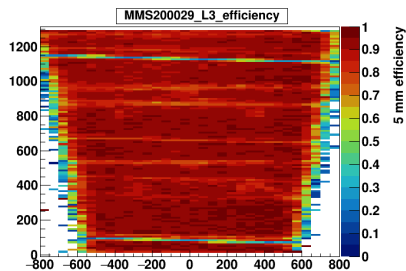


Comparison of External and Internal Efficiency Estimation (e.g. M29, stereo-in)

MDT tracks



internal



Theoretical Gain Description

Gain by Townsend (1910) :

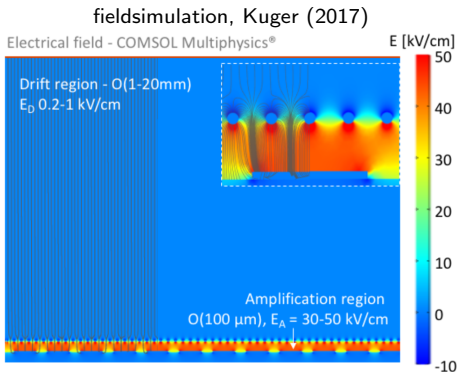
$$G = \exp [d \cdot \alpha]$$

α by Rose and Korff (1941) :

$$G = \exp \left[d \cdot \frac{pA}{T} \cdot \exp \left(-\frac{dpB}{TU} \right) \right]$$

effective description :

$$G = f(dp; U)$$

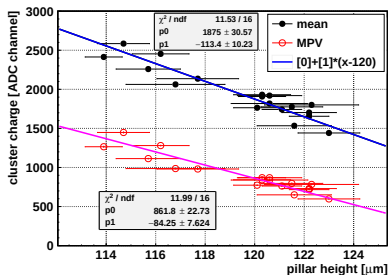


paper by Giomataris mentioning the prospects of a large sample of gap widths ($d \hat{=}$ pillar heights):

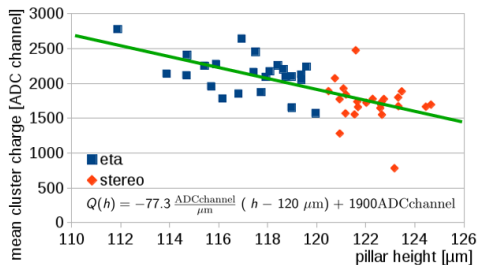
<https://www.slac.stanford.edu/pubs/icfa/fall99/paper1/paper1.pdf>

Pulse-Height Amplification-Gap Dependence

M8 and Eta3-Doublet
(same measurement period)



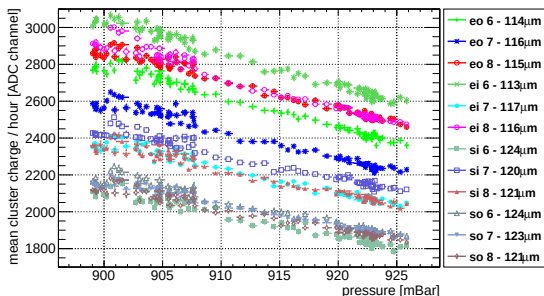
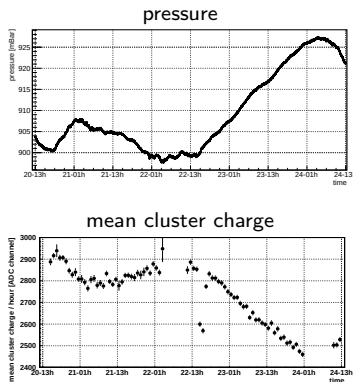
modules 1,3,6,7



$\Rightarrow$ for a single measurement period a good correlation is given

$\Rightarrow$ comparison of different periods require pressure compensation

Pulse-Height Pressure Dependence (M3 long term test)



⇒ clear pressure dependence, as expected

⇒ dependence on pillar-height NOT straight forward

(pressure biased: offset about 30 mBar)